



Fault detection and isolation in a photovoltaic system

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Abstract. This paper presents a fault detection and isolation method based on identification applied to a PV system. The process has been carried out by analyzing data from a real PV installation coupled with a grid-tied monophasic inverter.

Key words

Fault detection, photovoltaic system, Petri nets.

1. Introduction

Photovoltaic (PV) power generation has great importance as an alternative to conventional power sources; PV installations (in scales ranging from domestic uses to large solar farms) involves a significant investment that should be cared to preserve its safety and profitability. However a PV array can keep working in a fault state as a seeming normal behavior with power losses. In this case, when the fault state holds, permanent damages or premature aging in components could happen [1]. Therefore, PV fault diagnosis becomes important to performing fast maintenance and to improving power generation.

Most fault diagnosis techniques in PV systems are based on a model. These models usually start from electrical circuit equations as functions of irradiance and temperature to create a single solar cell, solar module model. The goal is obtaining an electric characterization [2], [3].

Some improvements on these classical analytical models increase their accuracy as: in [4] the authors consider miscellaneous power losses; in [5] the authors use weather satellite and ground weather stations data to feed the model simulations. Once obtained the model, real and simulated data are compared and the differences are analyzed. Another methods measure the PV system in normal conditions, after creates a real fault event and

analyses its effects; this information is used by [6], [7], [8], [9] to set up correlations between a known fault and the output signals, which are the basis for fault diagnosis. In [10], [11], [12], [13] the information is used in artificial intelligence or computational tools; these methods have reported high precision to detect and insulate fault events previously learned, but they need a training.

Regarding to large scale PV farms, in [14] the authors uses a Time Domain Reflectometry technique; a step voltage is introduced in a test electrical line, as a radar, the response is monitored by an oscilloscope to know type and localization fault. In [15], [16] use thermography and digital image processing to find hot spots and aging checking; these methods are accuracy but need fieldwork to recollect, analyze or record information, being less practical in large PV farms.

The methods mentioned above need wide technology knowledge and they can only diagnose faults included in the model. This work proposes a fault diagnosis method that works without fault models, so it could detect any abnormal behavior.

The main idea is identifying a normal working model of the PV system to compare real time data with the evolution of the identified model. Therefore, any deviation between prediction and observation implies necessarily a fault in the system.

The identified model is accurate and complete because the identification algorithm includes normal behavior signals and statistical stopping criteria to model stochastic processes.

The paper is organized as follows: Section 2 describes the fault detection and isolation process; in section 3 is

The part of the st-DICPN (Fig. 1) in green colour represents the normal behavior which is identified on-line from the observed legal sequences. The part of the net in black colour represents the identified fault behavior by means of applying the diagnostic algorithm. Arcs with black dashed lines are transformed into continuous lines when a fault is detected in the previous place and a fault transition fire in fault mode.

3. PV System fault diagnosis

The PV system is a real running installation coupled with a grid-tied monophasic inverter, extra information as solar radiance and surface temperatures, is provided by a weather sensor. **¡Error! No se encuentra el origen de la referencia.** shows the PV generator scheme and the signals to be considered in this work.

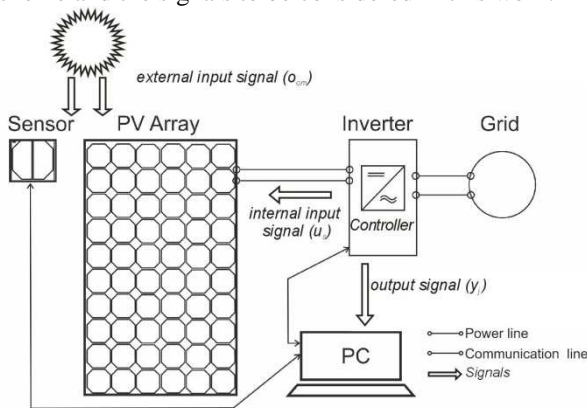


Fig. 2. PV Generator Scheme - power and communications.

Signals to consider are: external input signal: the solar radiation (W/m²); internal input signal: reference voltage (V); the system outputs: DC power (W), current (A), and cell surface temperature (°C).

¡Error! No se encuentra el origen de la referencia. shows electrical characteristics of the PV array.

Table I. – PV Array electric characteristics

Short Circuit Intensity (I_{sc})	4,85 A	Intensity Maximum (I_{Max})	4,39 A
Open Circuit Voltage (V_{oc})	445,2 V	Voltage Maximum (V_{Max})	354 V
Power Maximun (P_{Max})	2,500 kW		

Identification and monitoring data are read each 100 ms and the instant values are saved each one minute.

The system identification algorithm [17] works on-line and it requires binary signals.

Because the system works with continuous signals, a thresholding is needed to get the values. This process establishes operating ranges according to prior knowledge of the system performance. It is shown in **¡Error! No se encuentra el origen de la referencia.**

An event from the PV system is defined as $\omega^i = (u_s y_j) t^{ev}$ for the signals considered in **¡Error! No se encuentra el origen de la referencia.**, where u_s is a binary vector that represents the reference voltage range states; y_j is a binary vector representing the state of each range slot of the output signals.

For example, if the reference voltage is between 360 – 599, then the second range slot is active ([00010]). This binary array represents the binary number of 2; its event representation will be u_2 . If the DC power is between 1.200 – 1.800 watts, [0010], the current is between 0,002 – 0,6 A, [0010] and the cell surface temperature is between 0 – 16 °C, [1000], then, concatenating the input signals, the input vector is [001001001000]. This binary number is 584, therefore its event representation will be y_{584} . The input/output event will be $\omega^i = (u_2 y_{584}) \cdot t^{ev}$; being t^{ev} the elapsed time between ω^{i-1} and ω^i .

Table II. – Operating Ranges

Signal	Range
Solar Radiation (W/m ²)	>900
	200 – 900
	20 – 200
	0 – 20
Reference Voltage (V)	≥ 600
	360 – 599
	300 – 360
	180 – 300
DC Power (W)	0 – 180
	1.800 – 2.500
	1.200 – 1.800
	600 – 1.200
Current (A)	0 – 600
	3,8 – 4,39
	0,6 – 3,8
	0,002 – 0,6
Cell Surface Temperature (°C)	0 – 0,002
	>60
	25 – 60
	16 – 25
	0 – 16

The solar radiation is an external input signal. This signal defines an operating mode in the PV, o_{om} . Operation modes also needs a thresholding. As an example, $o_2 = [0100]$ means that the solar radiation is between 200 – 900 (W/m²).

We have feeded the identification algorithm with fault free data feeding and the result is a st-IPN representing the normal behavior of the system. **¡Error! No se encuentra el origen de la referencia.** shows the st-IPN structure.²

¡Error! No se encuentra el origen de la referencia. shows the st-IPN identified. Columns show (in order): transition number, input required, input place, initial reading, output place, expected reading and expected

² dY is omitted to improve the readability.

probability distribution for the time the transition fires. For example, first line defines the transition 1, when this transition is triggered by an u_1 input signal, the system goes from state 0 which has an output signal y_{1160} to state 1 which has an output signal y_{1096} . This transition fires with an average time of 17,5 min. if solar radiation is over

900 (W/m²) (o_1) or 56 min. if solar radiation is between 200 – 900 (W/m²) (o_2).

Table III. Structure st-IPN

Transition Nr	Input Function	pre		post		Density Function			
		state	Output Function	state	Output Function	o_1	o_2	o_4	o_8
1	u_1	0	y_{1160}	1	y_{1096}	N(17,5;1,21)	N(56,1;5,2)		
2	u_1	1	y_{1096}	2	y_{1096}	N(5;3,03)			
3	u_2	2	y_{1096}	3	y_{1160}	N(70,8;8,1)	N(4,25;1,25)		
4	u_2	3	y_{1160}	0	y_{1160}		N(7;1,03)		
5	u_4	0	y_{1160}	1	y_{1096}		N(1;0,3)		
6	u_4	3	y_{1160}	1	y_{1096}		N(3,75;0,82)		
7	u_2	1	y_{1096}	2	y_{1096}		N(3,6;0,35)		
8	u_4	2	y_{1096}	4	y_{584}		N(7,5;0,15)		
9	u_4	4	y_{584}	5	y_{552}		N(3,4;0,13)		
10	u_4	5	y_{552}	6	y_{548}		N(9;1,25)	N(8;2,05)	
11	u_4	6	y_{548}	7	y_{552}		N(12;1,3)	N(48;5,12)	
12	u_4	7	y_{552}	6	y_{548}		N(8,2;2,3)	N(116;5,8)	
13	u_4	6	y_{548}	8	y_{546}			N(24;3,01)	
14	u_4	8	y_{546}	9	y_{548}			N(3;0,325)	
15	u_4	9	y_{548}	10	y_{530}			N(28;2,04)	N(10;2,1)
16	u_4	10	y_{530}	11	y_{548}			N(20;2,14)	
17	u_4	11	y_{548}	7	y_{552}			N(42;2,16)	N(11;1,3)
18	u_4	7	y_{552}	12	y_{584}			N(3,2;0,2)	
19	u_4	12	y_{584}	5	y_{552}		N(63;2,03)	N(7,5;0,18)	
20	u_4	5	y_{552}	13	y_{552}		N(42;3,04)	N(98;4,06)	
21	u_2	13	y_{552}	6	y_{548}		N(5;0,2)	N(3;0,15)	
22	u_4	8	y_{546}	14	y_{530}	N(11;1)		N(3;0,5)	
23	u_4	14	y_{530}	15	y_{546}			N(7;0,01)	
24	u_4	15	y_{546}	14	y_{530}		N(9;0,1)	N(13,5;0,2)	
25	u_4	15	y_{546}	9	y_{548}		N(6,5;0,2)	N(8,5;0,3)	
26	u_4	9	y_{548}	8	y_{546}		N(4;0,1)	N(23;0,02)	
27	u_4	9	y_{548}	7	y_{552}			N(2;0,93)	
28	u_4	5	y_{552}	12	y_{584}			N(3,33;0,25)	
29	u_4	12	y_{584}	16	y_{1096}			N(36,75;3,1)	N(36;0,1)
30	u_4	16	y_{1096}	3	y_{1160}			N(15;1,9)	
31	u_4	3	y_{1160}	0	y_{1160}			N(1;0,6)	
32	u_2	0	y_{1160}	0	y_{1160}			N(23;2,82)	
33	u_1	1	y_{1096}	3	y_{1160}			N(3;0,13)	
34	u_1	3	y_{1160}	1	y_{1096}	N(4;0,2)		N(5;0,55)	
35	u_1	2	y_{1096}	2	y_{1096}	N(3,9;0,1)	N(17,3;4,1)		
36	u_4	2	y_{1096}	2	y_{1096}		N(44;5,5)		
37	u_2	1	y_{1096}	3	y_{1160}		N(63;2,5)		
38	u_2	2	y_{1096}	17	y_{1064}		N(23,8;2,03)		
39	u_4	17	y_{1064}	18	y_{552}		N(52;1)		
40	u_4	18	y_{552}	6	y_{584}		N(36;2,15)		
41	u_4	6	y_{584}	19	y_{292}		N(17;3,01)		
42	u_4	19	y_{292}	20	y_{548}			N(24,8;2,05)	
43	u_4	20	y_{548}	8	y_{546}			N(36;4,25)	N(10;2,2)
44	u_4	15	y_{546}	21	y_{290}			N(25,1;2,5)	
45	u_4	21	y_{290}	22	y_{274}			N(33;5,01)	
46	u_4	22	y_{274}	23	y_{530}			N(14,9;2,2)	
47	u_4	23	y_{530}	15	y_{546}		N(10;1,14)		
48	u_4	12	y_{584}	24	y_{648}			N(52,2;2)	
49	u_4	24	y_{648}	25	y_{1160}			N(5,65;0,97)	
50	u_2	25	y_{1160}	0	y_{1160}		N(75,8;4,12)		

This st-IPN structure is identifies three system behaviors; first, the system startup, until it is stabilizes; second, the

system working in maximum production and finally the power reduction.

The system startup is shown in Fig 2, as a st-IPN. We shows its working with the evolution of the first transition. The initial state is p_0 with $\varphi(p_0) = y_{1160}/[000000000000]$; $dY = \vec{0}$ because there is not history information of the output signals. If the solar radiation is between 20 – 200 W/m² (0_2) and after average time of 56,1 min. the input changes to u_1 and the net evolves to state p_1 with $\varphi(p_1) = y_{1096}/[0000 - 11000000]$, that means that the DC power continues between 600 – 1.200 W, the current changes its range and the cell surface temperature continues between 0 – 16 °C. But if the solar radiation is >900 (0_1), the input changes to u_4 and the net evolves to state p_1 , too, after average time 17,5 min.

Fig. 3. st-IPN.

A. On-line Monitoring

Fault detection and isolation consists of on-line monitoring; more precisely the process consists of:

- Initialize the diagnoser in (p_0), with $\varphi(p_0) = y_{1160}/\vec{0}$.
- Observe an on-line event, ω^i : if $\omega^i \in \mathcal{L}^N$, then $tr_r \in TR_N$ is fires in normal mode and a normal token is placed at $m(p_q) = \langle N \rangle$; else, a fault trace has been detected and a new $tr_r \in TR_F$ is generated (if not exist) and it is fires in f_{df} mode, a generic fault token is reached in $m(p_{VF}) = \langle (l, q)gf_{gf} \rangle$ and other integer token is reached in $m(p_{IF}) = \langle Integer \rangle$. The fault has been isolated.
- Observe an on-line new event ω^i .

Applying this process to the PV system, the results are shown in **!Error! No se encuentra el origen de la referencia.**

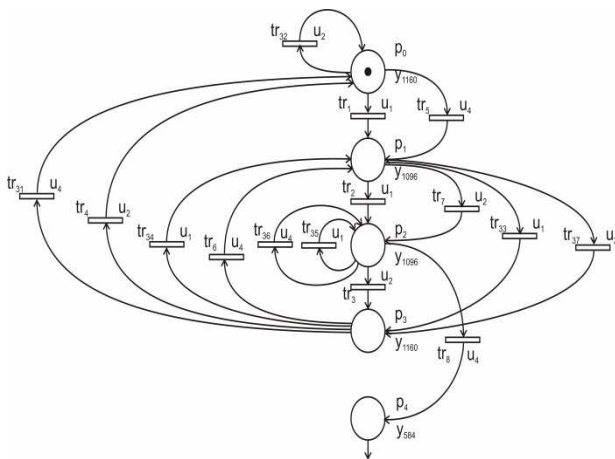


Table IV. Fault Detection and Isolation

Fault	Detection Time (min)	Fault Signal	State	Expected reading	Observed Reading	Faulty signal
f_1	153	Output	4	[001000101000]	[000000101000]	Power
f_2	304	Input	7	[00100]	[01000]	
f_3	307	Output	7	[001000100100] or [001001001000]	[001010001000]	Current
f_4	586	Output	12	[001000101000] or [010001001000] or [001001001000]	[000001001000]	Power
f_5	645	Output	24	[010010001000]	[001001001000]	Power and current
f_6	650	Input	24	[00100]	[01000]	
f_7	652	Input	24	[00100]	[01000]	
f_8	712	Output	24	[010010001000]	[010001001000]	Current

Table IV summarizes the principal faults detected in the on-line monitoring of the PV system carried out by four weeks.

The detected faults are mainly due to breakdown in communication by power shutdown, line insulation or by saturation of memory buffer.

B. Fault detection and isolation.

The diagnoser finds events that are related with an unwanted behavior. They occur when the system signals are not the expected or they do not occur at the expected time.

Faults f_1 and f_4 , are due to power readings outside the range; this mean that the power measurement is greater than 2.500 watts. This is not physically possible due to constraints of the inverter and can be inferred as a communication problem.

f_2 , is a fault shown a change in reference voltage without previous power change, then it is a bad control order creates by the inverter.

Faults f_3 , f_8 , are faults related to lower current, that could happen by a miscommunication or by a delay in device registers updating.

Fault f_5 is due to current and power signals fault. Two possible situations explains this. First, these behaviors are due to miscommunication in data acquiring. Second, the presence of partial shading that reduces solar radiation received by the photovoltaic generator but no in the radiation sensor or the other way round, resulting in a higher or lower power output vs. the expected power.

Faults f_6 and f_7 are faults related with voltage, they happen when there are suddenly changes in solar radiation and the Maximum Power Point Tracker algorithm oscillates for searching an optimum voltage operation. f_6 and f_7 are not a real faults, are actions to recover the normal operation not detected in the identification process.

The maximum number of faults that can be detected with this method depends on the number of signals to be measured in the system states. Therefore $|f_{df}| = 2^{n_{sr}}$; where n_{sr} is the number of sensors.

4. Conclusions

The detection and isolation faults method presented in this paper, applied in PV systems, has advantages with respect to classical techniques, because it works without previous behavior mode. It allows to detect all kind of faults starting from to the measures of the system signals based on residuals, over deterministic models. Moreover, the computational cost is low.

The "st-DICPN" proposed detects faults by comparing the behavior observed with normal language previously identified; therefore, the ability to detect faults strongly depends on the identified model.

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